

If it happens that the ratios of the crystal gram molecular refractivities are nearly equal to the corresponding ratios of the mean squares of the direction cosines of one of the molecular principal axes, two of the relative molecular refractivity uncertainties, those along the two other molecular principal axes, will be relatively large, e.g. if $R_a:R_b:R_c \simeq \alpha_1^2:\beta_1^2:\gamma_1^2$, then, as seen from equation (5), the determinants D_2 and D_3 will be very small, and conse-

quently, from equation (8), $\delta(r_M)/r_M$ and $\delta(r_N)/r_N$ will have rather high values.

Reference

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A new analytic approximation to atomic incoherent X-ray scattering intensities. By VEDENE H. SMITH JR, AJIT J. THAKKAR and DOUGLAS C. CHAPMAN, *Department of Chemistry, Queen's University, Kingston, Ontario K7L 3N6, Canada*

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A new analytic approximation to atomic incoherent X-ray scattering intensities is proposed. Unlike other approximations in the literature, the present function has the correct asymptotic behaviour at both large and small values of s . Fits to the incoherent intensities calculated by Cromer are presented for all atoms from He through Am.

Several analytic approximations to the intensity of the incoherent scattering of X-rays by N -electron atoms are available in the literature (Furukawa, Orton, Hamor & Williams, 1963; Rodriguez & Pings, 1965; Hajdu, 1971, 1972; Palinkas, 1972).

The more accurate approximations are those of Hajdu (1971, 1972) and of Palinkas (1972). Hajdu's function is given by

$$I_{\text{inc}}(s) = [Z - I_{\text{coh}}(s)Z^{-1}] [1 - M \exp(-Ks) + M \exp(-Ls)] \quad (1)$$

where Z is the atomic number, $s = \lambda^{-1} \sin \theta$, θ is twice the scattering angle, $2\pi/\lambda$ is the propagation constant of the X-ray, and

$$I_{\text{coh}}(s) = [F(s)]^2 \quad (2)$$

is the intensity of the coherent scattering where the form factor, $F(s)$, is approximated by

$$F(s) = \sum_{i=1}^4 A_i \exp(-B_i s^2) + C \quad (3)$$

Palinkas (1972) uses the following approximation

$$I_{\text{inc}}(s) = Z[1 - a(1 + bs)^{-c}] \quad (4)$$

Neither of the above functions has the correct asymptotic behavior either for large or for small values of s . The small- s behaviour of the total scattered intensity is known to be (Benesch & Smith, 1973)

$$I_{\text{tot}}(\mu) = N^2 - \mu^2 \langle r_{12}^2 \rangle / 3 + O(\mu^4) \quad (5)$$

where $\mu = 4\pi s$, and the small- s behaviour of the form factor is known to be (Benesch & Smith, 1973)

$$F(\mu) = N - \mu^2 \langle r^2 \rangle / 6 + O(\mu^4) \quad (6)$$

Hence the small- s behaviour of the incoherent scattering function is

$$I_{\text{inc}}(s) = a_2 s^2 + a_4 s^4 + \dots \quad (7)$$

where $a_2 = 16\pi^2(N\langle r^2 \rangle - \langle r_{12}^2 \rangle) / 3$. The large- s behaviour of the incoherent scattering function has been shown (Smith, 1970) to be

$$I_{\text{inc}}(s) = N + b_4 s^{-4} + b_6 s^{-6} + \dots \quad (8)$$

In this work we propose a new analytic approximation to the incoherent scattering function:

$$I_{\text{inc}}(s) = N[1 - (1 + as^2 + bs^4)(1 + cs^2 + ds^4)^{-2}] \quad (9)$$

Of course, for neutral atoms $N = Z$. As can be easily verified, equation (9) has the correct asymptotic behaviour for both large and small values of s .

Non-linear least-squares fits, over the range $0 \leq s \leq 8.0$, of equation (9) to the $I_{\text{inc}}(s)$ values calculated by Cromer & Mann (1967) and Cromer (1969) have been carried out with the algorithm of Fletcher (1971). Table 1 lists values of the parameters a, b, c , and d for all the atoms from He through Am. A measure of the 'goodness-of-fit', ε , is also listed in Table 1 for each atom. It is defined by

$$\varepsilon = 100 \left[\sum_{i=1}^k \delta_i^2 / (k-1) \right]^{1/2} / N \quad (10)$$

where δ_i are the deviations. ε can be considered to be a standard percent deviation because $N = \lim_{s \rightarrow \infty} I_{\text{inc}}(s)$. Table 1

shows that ε does not exceed $\sim 2\%$ for any of the atoms except Li. It should be noted that global optimization is currently impossible (Powell, 1970) and hence the parameters listed in Table 1 cannot be considered as defining anything better than a local minimum. Table 2 shows the typical quality of the fits for atoms of low ($Z=7$), medium ($Z=46$), and high ($Z=88$) atomic numbers.

A perusal of Table 1 shows that the fits worsen with increasing atomic number. For atoms of low and medium atomic number the overall quality of the fits obtained is comparable to the fits of Hajdu (1972) and Palinkas (1972). For atoms of high atomic number the fits are somewhat inferior to those of Palinkas (1972). However equation (9)

Table 1. Parameters for the analytic approximation, equation (9), to incoherent scattered X-ray intensities; the notation used is: $1.5566D \text{ O}1 = 1.5566 \times 10^1$.

| Atom Z | a | b | c | d | ϵ |
|--------|----|------------|-------------|------------|------------|
| HE | 2 | 7.2391D-01 | -2.1464D-01 | 5.1019D 00 | 1.5566D C1 |
| LI | 3 | 2.6676D 01 | 6.8617D 00 | 2.6259D 01 | 0.0112D C1 |
| BE | 4 | 1.6566D 01 | 1.1177C 03 | 4.0948D 01 | 1.0399D C2 |
| B | 5 | 7.2557C 00 | 2.7281D 02 | 2.2653D C1 | 3.4403D C1 |
| NO | 6 | 4.3535C 01 | 9.3125C 01 | 1.4654D 01 | 1.4628D C1 |
| O | 7 | 4.5051D 00 | 4.0565D 01 | 1.1061D 01 | 1.3225D C0 |
| N | 8 | 3.2434C 00 | 1.9377C 01 | 8.2735D C0 | 4.0087D C0 |
| NE | 10 | 3.0545D 00 | 1.0031D 01 | 6.5952D 00 | 2.3142D C0 |
| NA | 11 | 1.6255D 01 | 4.5169D 01 | 1.3167D 01 | 5.4559D C0 |
| MG | 12 | 1.6542D 02 | 2.7409C 03 | 7.0762D 01 | 7.7161C 01 |
| AL | 13 | 1.6542D 02 | 2.7409C 03 | 7.0762D 01 | 7.7161C 01 |
| SI | 14 | 8.0715C 01 | 1.5284C 03 | 5.2054D 01 | 3.6858C 01 |
| S | 16 | 5.9522D 01 | 7.0348D 02 | 3.7402D 01 | 2.0872D 01 |
| CL | 17 | 3.4552C 01 | 2.5266C 02 | 2.4481D 01 | 1.8430D C0 |
| AR | 18 | 7.1232D C1 | 1.5569D 02 | 1.5657D 01 | 5.8774D C0 |
| K | 19 | 1.9276D 01 | 7.7510D 01 | 1.3788D 01 | 5.1378D C0 |
| CA | 20 | 2.6545D 01 | 2.8598D 02 | 2.5590C 01 | 6.5619D C0 |
| SC | 22 | 3.3212D 01 | 2.3259D 02 | 2.3472D C1 | 6.8604D C0 |
| TI | 23 | 1.8262C 01 | 1.8205C 02 | 2.0962D 01 | 5.7117D 00 |
| V | 23 | 6.6153D 01 | 1.4376C 03 | 1.8283D 01 | 4.8164D C0 |
| CR | 24 | 1.8529D 01 | 7.2508C 01 | 1.3819D 01 | 3.0415D C0 |
| MN | 25 | 2.1529D 01 | 4.4848C 01 | 1.5188D 01 | 3.6264D C0 |
| FE | 26 | 1.7616D 01 | 6.4657D 01 | 1.2621D 01 | 7.7056D C0 |
| CO | 27 | 1.6346D C1 | 5.5896D 01 | 1.1687C 01 | 2.4272D C0 |
| CU | 28 | 1.5151D 01 | 4.4546D 01 | 1.3436D C0 | 1.7194D C0 |
| ZN | 30 | 1.4619C 01 | 4.4546D 01 | 1.1836D 01 | 2.5707C 01 |
| GA | 31 | 1.7176D C1 | 6.4652D 01 | 1.1820D 01 | 2.9211D C0 |
| AS | 32 | 1.7471D 01 | 8.2421C 01 | 1.3080D 01 | 2.9211D C0 |
| SE | 33 | 1.7471D 01 | 8.2421C 01 | 1.3080D 01 | 2.9211D C0 |
| KR | 34 | 2.0255D 01 | 5.2597D 01 | 1.3650D 01 | 2.8488D C0 |
| BR | 35 | 2.0211C 01 | 9.2524C 01 | 1.3623D C1 | 2.7282D C0 |
| R | 37 | 1.2650D 01 | 1.0161D 02 | 1.4409D 01 | 3.3584D C0 |
| Y | 37 | 1.4449D C1 | 1.3593C 02 | 1.6192D 01 | 3.3017D C0 |
| Zr | 40 | 2.3652D 01 | 1.2653D 02 | 1.6237D 01 | 2.0148D C0 |
| Nb | 41 | 1.9178C 01 | 8.1654D 01 | 1.2952D 01 | 1.5678D C0 |
| Mo | 42 | 1.6151D 01 | 6.7437C 01 | 1.1867D 01 | 1.6597C 01 |
| Tc | 43 | 1.8262C 01 | 7.3766C 01 | 1.2323D 01 | 1.5131D C0 |
| Ru | 44 | 1.5421D 01 | 5.1534C 01 | 1.0601D 01 | 1.3610D C0 |
| Rh | 45 | 1.8262C 01 | 4.3514C 01 | 9.8388D C0 | 1.1954D C0 |
| Pd | 46 | 1.8262C 01 | 4.3514C 01 | 9.8388D C0 | 1.1954D C0 |
| Ag | 47 | 1.1510D 01 | 3.0295C 01 | 8.3388D 00 | 9.0612D C1 |
| Cd | 48 | 1.8262D 01 | 3.2428D 01 | 8.5514D 00 | 5.3212D C1 |
| In | 49 | 1.8262D 01 | 3.2428D 01 | 8.5514D 00 | 5.3212D C1 |
| Sn | 50 | 1.3547D 01 | 4.0031D 01 | 9.3037D 00 | 1.0356C 01 |
| Sb | 51 | 1.3466D 01 | 3.9457D 01 | 9.2631D 00 | 1.0057D C0 |
| Te | 52 | 1.4450D 01 | 4.6190C 01 | 9.8671C 00 | 1.0869D C0 |
| I | 53 | 1.4575D 01 | 4.9456D 01 | 1.0161D 01 | 1.1116C 01 |
| Xe | 54 | 1.4888D 01 | 4.8883C 01 | 1.0102D 01 | 1.0775D C0 |
| Cs | 55 | 1.6151D 01 | 5.7596D 01 | 1.0886D 01 | 1.1773D C0 |
| Ba | 56 | 1.7675D 01 | 1.1185D 01 | 1.1185D 01 | 1.1185D C0 |
| La | 57 | 1.8262D 01 | 6.4363D 01 | 1.2167D 01 | 1.3151D C0 |
| Ce | 58 | 1.8262D 01 | 6.4363D 01 | 1.2167D 01 | 1.3151D C0 |
| Pr | 59 | 1.8262D 01 | 6.4363D 01 | 1.2167D 01 | 1.3151D C0 |
| Nd | 60 | 1.6148D 01 | 5.8598C 01 | 1.0747C 01 | 1.1155D C0 |
| Pm | 61 | 1.5741D 01 | 5.5777D C1 | 1.0458D 01 | 1.0774D C0 |
| Sm | 62 | 1.8262D 01 | 6.4363D 01 | 1.2167D 01 | 1.3151D C0 |
| Eu | 63 | 1.5057C 01 | 5.1053C 01 | 9.9393D 00 | 1.0356D C0 |
| Gd | 64 | 1.4648D 01 | 5.0133D 01 | 9.8426D 00 | 9.7728D C1 |
| Tb | 65 | 1.4484D 01 | 4.7248C 01 | 9.5464D 00 | 9.3253D C1 |
| Dy | 66 | 1.3755C 01 | 4.3137C 01 | 8.8604D 00 | 8.4486D C1 |
| Ho | 67 | 1.3447D 01 | 4.1045D 01 | 8.6467D C0 | 8.1464D C1 |
| Er | 68 | 1.3146C 01 | 3.9287C 01 | 8.4562C 01 | 7.9066C 01 |
| Yb | 69 | 1.2822D 01 | 3.7603D 01 | 8.2568D C0 | 7.6627D C1 |
| Tm | 70 | 1.2667C 01 | 3.6643D 01 | 8.2264C 00 | 7.4720D C1 |
| Lu | 71 | 1.2576D 01 | 3.6129C 01 | 8.1062D 00 | 7.2092D C1 |
| HF | 72 | 1.2532D 01 | 3.5877C 01 | 8.0420D 00 | 6.7752D C1 |
| Ta | 73 | 1.1566D 01 | 2.9720C 01 | 7.5123C 00 | 6.2773D C1 |
| W | 74 | 1.1459D 01 | 2.8660D 01 | 7.1515D 00 | 5.7682D C1 |
| Os | 75 | 1.0841D 01 | 2.6277C 01 | 6.5734D 00 | 4.6176D C1 |
| Ir | 76 | 1.0530D 01 | 2.5229C 01 | 6.4789D 00 | 4.5891D C1 |
| Pt | 77 | 1.0255D 01 | 2.3781D 01 | 5.8116D 00 | 4.3980D C1 |
| IR | 78 | 9.1444D 00 | 1.8784D 01 | 5.8116D 00 | 4.3980D C1 |
| Au | 79 | 8.6652D 00 | 1.6793C 01 | 5.8116D 00 | 4.3980D C1 |
| Hg | 80 | 8.1670D 00 | 1.7414D C1 | 5.5035D C0 | 4.0768D C1 |
| TL | 81 | 9.0767C 00 | 1.8476C 01 | 6.4519D C0 | 4.1721D C1 |
| Pb | 82 | 9.3622D 00 | 1.9739C 01 | 6.2852D C0 | 4.2863D C1 |
| Bi | 83 | 9.5574D 00 | 2.0605D 01 | 6.3414D 00 | 4.3371D C1 |
| PO | 84 | 9.2556D 00 | 2.2287C 01 | 6.5024D 00 | 4.4878D C1 |
| FR | 85 | 1.0267D 01 | 3.902C 01 | 6.7681D 00 | 4.6167D C1 |
| RA | 86 | 1.0425D 01 | 4.4633D 01 | 6.8643D 00 | 4.6342D C1 |
| FR | 87 | 1.1077D 01 | 7.5477C 01 | 7.2583D 00 | 4.9341D C1 |
| AC | 89 | 1.2372D 01 | 5.070D C1 | 7.760D C0 | 5.3275C 01 |
| TH | 90 | 1.2708D 01 | 7.029D C1 | 8.240D C0 | 5.6015D C1 |
| Pa | 91 | 1.1521D 01 | 3.2573C 01 | 7.7684D 00 | 5.1436D C1 |
| U | 92 | 1.1596D 01 | 3.0746C 01 | 7.589D 00 | 4.9185D C1 |
| NP | 93 | 1.1277D 01 | 3.9055D 01 | 7.3579C 00 | 4.7057D C1 |
| PU | 94 | 1.0255D 01 | 5.5479D 01 | 6.5123D C0 | 4.3325D C1 |
| AM | 95 | 1.0243D 01 | 5.3518D 01 | 6.7044D C0 | 4.142D C1 |

is the only one that has the correct asymptotic behaviour at large and small values of s . Hence the present fits should be especially useful in studies of electron scattering where accuracy of the fits for $s \leq 0.1$ is rather important.

Table 2. Comparison of incoherent scattered X-ray intensities computed by Cromer, I_e , with those predicted by equation (9), I_f .

| | s | I_e | I_f | I_e | I_f | I_e | I_f |
|--|-------|-------|-------|--------|--------|--------|--------|
| | | Z=7 | | Z=46 | | Z=88 | |
| | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 |
| | 0.005 | 0.003 | 0.003 | 0.010 | 0.006 | 0.036 | 0.008 |
| | 0.010 | 0.013 | 0.012 | 0.039 | 0.022 | 0.144 | 0.032 |
| | 0.050 | 0.310 | 0.296 | 0.923 | 0.548 | 2.520 | 0.780 |
| | 0.100 | 1.080 | 1.054 | 3.103 | 2.068 | 5.696 | 2.958 |
| | 0.150 | 2.003 | 1.994 | 5.536 | 4.251 | 8.858 | 6.122 |
| | 0.200 | 2.858 | 2.876 | 7.725 | 6.747 | 11.964 | 9.789 |
| | 0.300 | 4.097 | 4.120 | 11.441 | 11.646 | 17.672 | 17.152 |
| | 0.400 | 4.792 | 4.783 | 14.824 | 15.685 | 22.270 | 23.378 |
| | 0.500 | 5.182 | 5.160 | 17.943 | 18.786 | 26.283 | 28.247 |
| | 0.600 | 5.437 | 5.422 | 20.653 | 21.173 | 30.139 | 32.040 |
| | 0.700 | 5.635 | 5.635 | 22.904 | 23.071 | 33.790 | 35.080 |
| | 0.800 | 5.809 | 5.821 | 24.756 | 24.645 | 37.136 | 37.617 |
| | 0.900 | 5.968 | 5.988 | 26.316 | 26.002 | 40.121 | 39.823 |
| | 1.000 | 6.113 | 6.134 | 27.677 | 27.211 | 42.744 | 41.811 |
| | 1.500 | 6.630 | 6.614 | 32.888 | 32.104 | 52.502 | 50.271 |
| | 2.000 | 6.860 | 6.821 | 36.349 | 35.881 | 59.860 | 57.658 |
| | 3.000 | 6.979 | 6.951 | 40.389 | 40.848 | 69.885 | 69.484 |
| | 4.000 | 6.996 | 6.982 | 42.529 | 43.357 | 75.363 | 77.077 |
| | 5.000 | 6.999 | 6.992 | 43.658 | 44.579 | 78.664 | 81.509 |
| | 8.000 | 7.000 | 6.999 | 45.019 | 45.694 | 83.881 | 86.363 |

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